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The Future Photon Initiative

CfD Annual Report 2015

Welcome to the Center for Detectors. CfD designs, develops, and implements new advanced sensor technologies through collaboration with academic researchers, industry engineers, government scientists, and university/college students. The CfD operates four laboratories and has approximately a dozen funded projects to advance detectors in a broad array of applications, e.g. astrophysics, biomedical imaging, Earth system science, and inter-planetary travel. Our observational astrophysics programs include studies of massive stars, massive star clusters, the Galactic center, the interstellar medium, and the cosmology of structure formation in the Universe.



Student involvement is central to all CfD projects. To obtain a PhD degree with us, apply to [Astrophysical Sciences and Technology](#), [Microsystems Engineering](#), or [Imaging Science](#). To obtain an MS degree with us, contact the CfD Director.

[CfD feature article in RIT Research Magazine](#)

[School of Design Collaboration with CfD](#)



CfD News

February, 2016

[CfD Part of the Future Photon Initiative](#)

November, 2015

[CfD Wins NSF Funding to Develop Next Generation Infrared Detectors](#)

October, 2015

CfD Welcomes New Faculty Member Parsian Mohseni

[CfD Alumni are featured in the Winter 2015-2016 issue of University Magazine](#)

September, 2015

[CfD is featured in RIT's University News Campus Spotlight](#)

August, 2015

[CfD Faculty Member Stefan Preble's work is featured in the cover article of the September issue of Rochester Engineering Society magazine](#)

July, 2015

[CfD Faculty Member Michael Zemcov Part of Team Chosen to Develop Space Mission to Study Neutron Stars, Black Holes and More](#)

[CfD Welcomes New Faculty Member Jing Zhang](#)

[CfD Welcomes New Faculty Member Stefan Preble and His Nanophotonics Group](#)

[CfD to be part of New Rochester-based Photonics Center](#)

June, 2015

[CfD Welcomes New Faculty Member Michael Zemcov](#)

May, 2015

[RIT Convocation Speaker, NASA Administrator Maj. Gen. Charles F. Bolden Jr., Visits CfD, Pixieset](#)



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April, 2015

CfD Director Gives Invited Presentation at the Hartford Club

CfD Director's image of the Pistol Star featured in Jersey Post Stamps commemorating 25 years of the Hubble Space Telescope RIT News Jersey Post

NASA Astronaut Dr. Donald Pettit Visits CfD [Facebook](#) [Flickr](#) [RIT News](#)

March, 2015

[CfD Student Kim Kolb Defends PhD Thesis](#)

CfD Hosts Sloane Wiktorowicz's presentation "Characterizing Exoplanet Atmospheres through Our Own" [Announcement Video](#)

February, 2015

CfD Hosts Michael Zemcov's presentation "Mapping the Large Scale Structure of the Cosmos from the Big Bang to the Present" [Announcement Video](#)

CfD Hosts Roger O'Brient's presentation "Moore's Law and Astrophysics: Detectors that Teach Us About Physics Beyond the Standard Model" [Announcement Video](#)

September, 2014

[CfD Annual Report 2014](#)

August, 2014

[CfD Director appears on Connections: Science Roundtable with Evan Dawson on WXXI AM 1370 News](#)

April, 2014

[CfD Hosts Minh Nguyen's presentation "Nanoscale Semiconductor Hybrids: Fundamental Physics and Advanced Devices"](#)

February, 2014

[Infrared Detector Technology at RIT's Center for Detectors](#)

November, 2013

[CfD Shedding light on Earth-like planets](#)

October, 2013

[CfD Graduate Student Kimberly Kolb Studies Tiny Details of Vast Universe](#)

September, 2013

[CfD Annual Report 2013](#)



April, 2013

[CfD Ph.D. Gives Christine Trombley Momentum](#)

[CfD Hosts Christoph Baranec's Presentation "Rayleigh Laser Guide Stars Pioneering the Next Decade of Astronomical Adaptive Optics"](#)

[CfD Executive Assistant Allison Conte Helps Brand The Research Center](#)

March, 2013

CfD Ph.D. Student Christine Trombley wins the NASA/Cornell University Graduate Research Fellowship in Astrophysical Sciences and Technology

January, 2013

[NASA Awards CfD \\$1.1 million \(RBJ Article\)](#)

September, 2012

[CfD Annual Report 2012](#)

July, 2012

[CfD Leads Development of Next-generation Infrared Detectors](#)

[Detector Virtual Workshop: Hooman Mohseni presents "Nano-injection Detectors and Imagers" on 7/13/2012 at RIT](#)

June, 2012

[Secret Spy Telescopes Made in Rochester Video](#)

[CfD Student Brian Glod Completes Master's Degree](#)

[CfD Administrative Assistant Adena Thomson Thrives at RIT](#)

[Detector Virtual Workshop: Jim Bangs presents on Single and Dual Band HgCdTe Detectors on 6/4/2012 at RIT](#)

May, 2012

[Detector Virtual Workshop: Sanjay Krishna presents "The Infrared Retina: Ushering in the Fourth Generation of Infrared Detectors" on 5/14/2012](#)

April, 2012

[Detector Virtual Workshop: Michael Krainak presents "Candidate detectors for space-qualified time-resolved photon counting" on 4/23/2012](#)

March, 2012

[Detector Virtual Workshop: Jim Beletic presents "The Fantastical Discoveries of Astronomy made possible by the Wonderful Properties of II-VI Materials" on 3/26/2012 at](#)



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[Detector Virtual Workshop: Mark Itzler presents "Photon Counting with InGaAsP Single Photon Avalanche Diodes" on 3/12/2012 at RIT](#)

February, 2012

[Detector Virtual Workshop: Kieran O'Brien presents "MKIDs" on 2/27/2012 at RIT](#)

[Detector Virtual Workshop: Bernard Rauscher presents "Reducing Read Noise of JWST Detectors" on 2/6/2012](#)

December, 2011

[Detector Virtual Workshop: Craig Mackay presents "EMCCDs" on 12/5/2011](#)

November, 2011

[Detector Virtual Workshop: Alan Middall presents "Quantum Communication Detectors" on 11/28/2011 at RIT](#)

[Detector Virtual Workshop: Karl Berggren presents "Superconducting Single Photon Detectors" on 11/14/11 at RIT](#)

[Jon Morse presents "Frontiers in Astrophysics and the Federal Budget Landscape" Announcement](#) [Adobe Connect Video](#) [PDF](#)

[Mysterious Absorption Lines Could Illuminate 90-year Puzzle](#)

October, 2011

[Bahram Mobasher presents "Future Big Telescopes"](#)

[Detector Virtual Workshop: Andy Shearer presents "High Time Resolution Astrophysics"](#)

[CfD Annual Report 2011](#)

[Hans Zinnecker presents "SOFIA Infrared Observatory and its recent results"](#)

September, 2011

[CfD Initiates Detector Virtual Workshop speaker series](#)

[CfD Graduate Student Christine Trombley Wins Prestigious NASA Graduate Student Researchers Program Fellowship](#)

[Detector Virtual Workshop: Dr. Robert Hadfield presents "Infrared Superconducting Single Photon Detectors" Announcement](#) [Video](#)

[CfD Student Member Adena Thomson wins Accounting Scholarship](#)

August, 2011

[CfD Director Invited to Speak at Central Kiloparsec in Galactic Nuclei Conference](#)

[Powerpoint](#) [Adobe Connect Video](#)



CfD Director Invited to Speak at Single Photon Imaging Conference [Powerpoint PDF](#)

CfD Director Invited to Speak at Techniques and Instrumentation for Detection of Exoplanets Conference [Powerpoint PDF](#)

July, 2011

[CfD Graduate Student Kimberly Kolb Completes M.S. Thesis](#)

June, 2011

[CfD Director Invited to Speak at Single Photon Workshop 2011](#) [Adobe Connect Video](#)

May, 2011

Donald Hall Presents at Center for Detectors

[Announcement](#) [Adobe Connect Video](#) [Quicktime File](#)

April, 2011

[Center for Detectors Featured in Rochester Business Journal](#)

March, 2011

[School of Design collaborates with RIDL](#)

February, 2011

[Center for Detectors is Established](#)

[CfD Student Chris Maloney Reports on FiOS Conference](#)

December, 2010

Chris Packham Presents at Center for Detectors

[Announcement](#) [Powerpoint](#) [Quicktime File](#) [Adobe Connect Video](#)

[RIDL Director Invited to Speak at NASA/JPL](#)

October, 2010

Joss Hawthorn Presents at Center for Detectors

[Announcement](#) [WMV File](#) [Quicktime File](#)

June, 2010

RIDL Presents at SPIE, San Diego, CA

Kolb et al., Hybridization of a sigma-delta-based CMOS hybrid detector
[PDF](#) [Powerpoint](#)

Hanold et al., Characterization of a sigma-delta-based Monolithic CMOS Detector
[PDF](#) [Powerpoint](#)



May, 2010

RIDL Director Invited to Give NRO Technology Seminar

February, 2010

[RIDL Selected by NASA to Develop Photon Counting Imaging Detectors for Space Exoplanet Finding Missions](#)

January, 2010

RIDL Director Invited to Speak at Keck Institute for Space Studies, Caltech [\(PDF\)](#) [\(VIDEO\)](#)

October, 2009

[RIDL Director Invited to Speak at AstroMed09, Sydney, Australia](#)

[RIDL Director Invited to Speak at Detectors for Astronomy 2009, Munich, Germany](#)

September, 2009

[Youtube Video of RIDL on The History Channel \(part 1\) \(part 2\) \(part 3\)](#)

[RIDL Student Kim Kolb Awarded Prestigious BAE Fellowship \(Photo and caption\)](#)

[RIDL Featured on Rochester CBS/Fox News Channel 8 and 31 \(video\)](#)

July, 2009

[A Day in the Life of the RIDL Director in Astronomy Magazine](#)

[RIDL to be Featured on History Channel](#)

May, 2009

[RIDL Featured in RIT Research Magazine](#)

Bruce Tromberg Presents at Center for Detectors

[MP4 File](#) [Adobe Connect Video](#)

March, 2009

RIDL Hosts Quantum Limited Detector Workshop

[Press Release](#)

[RIT Presents: The Next Generation of Imaging Detectors](#)

[Agenda with Recorded Presentations and Slides](#)

[Workshop Web Site](#)

October, 2008

Moore Foundation Awards RIDL \$2.8 Million to Develop "Noiseless" Detector



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Moore Foundation Awards RIDL \$2.8 Million to Develop "Noiseless" Detector

Press Release

[Moore Foundation Awards RIT Scientist \\$2.8 Million to Develop "Noiseless" Detector](#)

News Articles

[Rochester Democrat and Chronicle](#)

[Laser Focus World](#)

Shouleh Nikzad Presents at Center for Detectors

[Announcement WMV File](#)

Michael Hoenk Presents at Center for Detectors

[Announcement WMV File](#)

April 2008

RIDL Receives NASA Planetary and Earth Sciences Detector Development Grant

Press Release

[LIDAR Imaging Detector Could Build 'Super Road Maps' of Planets and Moons](#)

News Articles

[Channel 10 News Video \(Associated article\)](#)

[MIT Technology Review](#)

[Optics.org](#)

[Astrobiology Magazine](#)

[United Press International](#)

[Popular Science](#)

[Spacemart.com](#)

[PHYSORG.com](#)

[Universe Today](#)

[Science Daily](#)

[Photonics.com](#)

[The Times of India](#)

[RIT University News](#)

Bedabrata Pain Presents at Center for Detectors

[WMV File](#)

August 2007

RIDL Receives NASA Planetary Science Detector Development Grant

Press Release

[New Imaging Detectors Could Take Snapshots from Deep Space](#)

News Articles

[MSNBC: New Cameras Can Withstand Space Radiation](#)



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[Press Release](#)

[LIDAR Imaging Detector Could Build 'Super Road Maps' of Planets and Moons](#)

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[Popular Science](#)

[Spacemart.com](#)

[PHYSORG.com](#)

[Universe Today](#)

[Science Daily](#)

[Photonics.com](#)

[The Times of India](#)

[RIT University News](#)

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August 2007

RIDL Receives NASA Planetary Science Detector Development Grant

[Press Release](#)

[New Imaging Detectors Could Take Snapshots from Deep Space](#)

News Articles

[MSNBC: New Cameras Can Withstand Space Radiation](#)

[Space.com: Tougher cameras Designed to Withstand Space Radiation](#)

[Photonics.com: Detector Resists Radiation](#)

May 2007

[RIDL Selected by NASA to Develop New Deep Space Detectors](#)

March 2007

[RIDL Achieves First Light with 4Kx4K SiPIN Detector](#)

October 2006

[RIDL Chosen to Lead the LSST Guider Project](#)

[RIDL Chosen to Perform SNAP Prototype Detector Testing](#)

February 2006

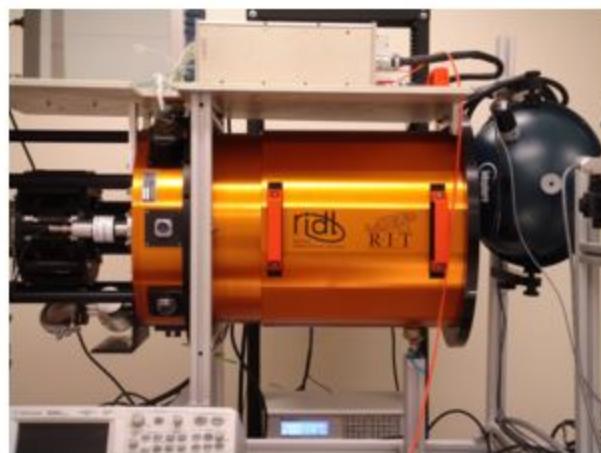
[RIDL Director Receives NYSTAR Faculty Development Award](#)

[Figer Massive Stars Group News](#)



Facilities, Capabilities, and Equipment at the Center for Detectors

The Center for Detectors (CfD) is located in the IT Collaboratory and has large contiguous spaces for offices and labs, including offices for 17 people, and four research laboratories: the Rochester Imaging Detector Laboratory, The Imaging LIDAR Laboratory, the Quantum Dot Detector Laboratory, and the Wafer Probe Station Laboratory. There are approximately 5000 square feet of space, divided equally between labs and offices. The laboratories contain special facilities and equipment dedicated to the development of detectors.



Dewar 1

The RIDL detector testing systems use cylindrical vacuum cryogenic dewars. Each individual system uses a cryo-cooler that has two cooling stages: one at ~ 10 K (7 W) and another at ~ 60 K (70 W). The cold temperatures yield lower detector dark current and read noise. The systems use Lakeshore Model 340 temperature controllers to sense temperatures at 10 locations within the dewars and control a heater in the detector thermal path. This thermal control system stabilizes the detector thermal block to 400 micro-K RMS over timescales greater than 24 hours.



Dewar 2

The detector readout systems include an Astronomical Research Cameras controller having 32 digitizing channels with 1 MHz readout speed and 16-bit readout capability, two Teledyne SIDECAR ASICs having 36 channels and readout speeds up to 5 MHz at 12-bits and 500 kHz at 16-bits, and custom FPGA systems based on Altera and Xilinx parts. The controllers drive signals through cable harnesses that interface with Detector Customization Circuits (DCCs), which are designed in-house and consist of multi-layer cryogenic flex boards. The DCCs terminate in a single connector, which then mate to the detector connector. Three-axis motorized stages provide automated lateral and piston target adjustment.





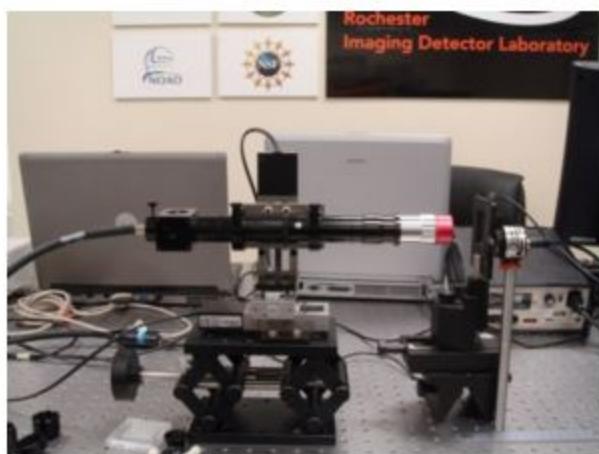
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Dewar 3

Two of the dewars have a side-looking port that is useful for exposing detectors to high energy radiation beams. The lab has a large integrating sphere that provides uniform and calibrated illumination from the ultraviolet to the infrared. The dewars are stationed on large optical tables that have vibration-isolation legs.



Spot projector with automated three axis motion control system

The Rochester Imaging Detector Laboratory has a Pico Quant pulsed laser for LIDAR system characterization and other testing that requires pulsed illumination.

RIDL also has monochromators with specialized light sources for illumination from the UV into the IR (250 nm – 2500 nm).

NIST-traceable calibrated photodiodes (with a wavelength range of 200 nm – 1100 nm) provide for absolute flux measurements.

RIDL also has a spot projector to characterize the inter-pixel response of the detectors, including optical and electrical crosstalk. The image to the left shows a laser spot projection system on a 3D motorized stage that produces a small point source for measurements of intrapixel sensitivity.



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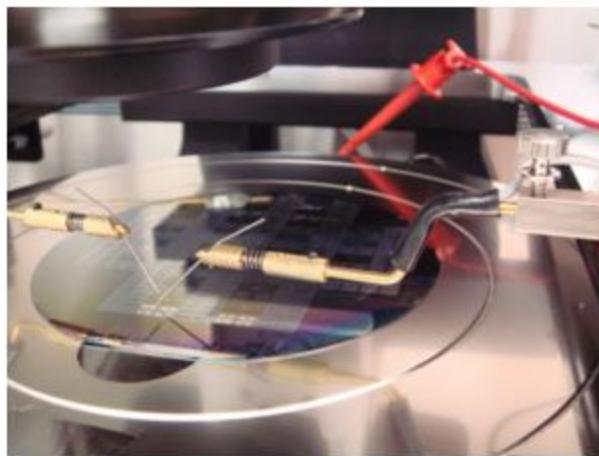
Lab Area in RIDL

The lab contains eight data reduction PCs each with eight processors and up to 16 GB memory for data acquisition, reduction, analysis and simulations and 25 TB of data storage. Custom software runs an automated detector test suite of experiments. The test suite accommodates a wide variety of testing parameters through the use of parameter files. A complete test suite takes a few weeks to execute and produces ~0.5 TB of data. The data reduction computers reduce and analyze the data using custom automated code, producing publication-quality plots in near-real time as the data are taken.



RIDL Clean Room

CfD has a dedicated class 1000 cleanroom (by FED Standard 209E), located in the [SMFL](#). The SMFL has 10,000 ft² of additional cleanroom space in class 1000, 100, and 10. Using the SMFL's resources, the Center can fabricate detectors with custom process flows and the freedom to use multiple process variations.



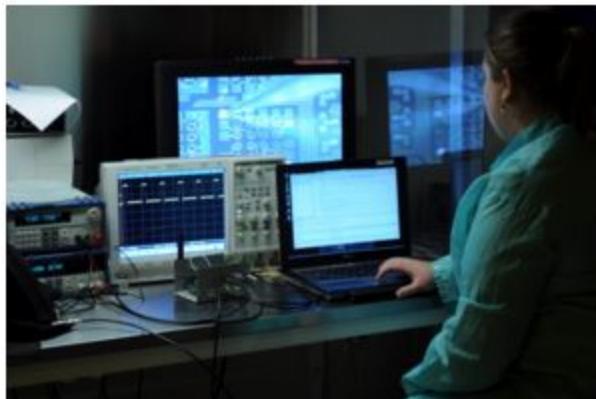
Testing a p-i-n diode device wafer using the probe station

The Center's cleanroom and probe stations allow wafer-level testing, even during the fabrication process, allowing mid-process design changes. The probe station accommodates electrical and circuit analysis of both wafers and packaged parts, including low current and radio frequency (RF) probing. Also available for the CfD's use are the Amray 1830 Scanning Electron Microscope (SEM), used for high-magnification imaging of devices, and the WYKO white light interferometer, used for surface topography measurements. The SMFL also has other in-line fabrication metrology capabilities, including material layer thickness, refractive index, and wafer stress characterization tools.



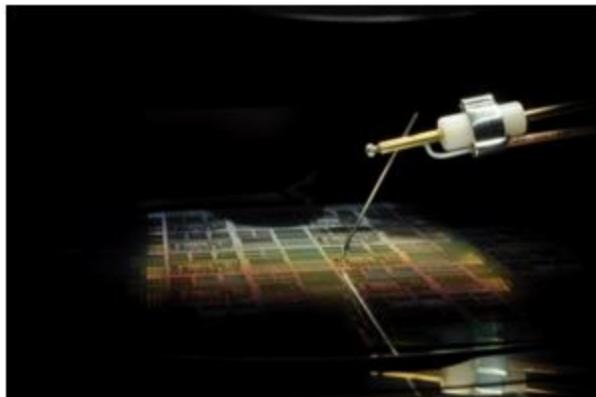
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PhD student Kimberly Kolb conducts electrical experiments on one of the cutting edge devices being characterized at the Center for Detectors.

The customized setup shown here is comprised of two voltage power supplies, an Agilent oscilloscope, an LCD screen for viewing the devices through the microscope probe station, and a custom circuit board for specific device diagnostics. The dedicated lab computer also runs a specially-designed data acquisition program to collect and analyze data from the device.



This image is a close-up of a device wafer being tested on the probe station under dark conditions.

The entire probe station is covered so that no stray light enters the testing environment. These conditions provide the basis for valuable testing and data analysis. The probe tip is contacting a single test device via a metal pad with dimensions of only 70 microns by 70 microns (an area of 0.0049 mm^2).

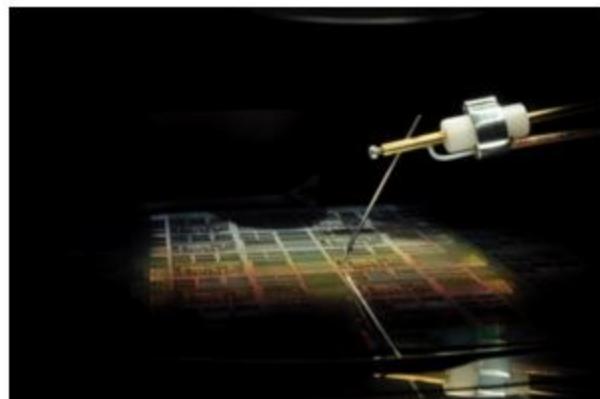


Many other RIT facilities are available for use by the Center for Detectors. The Brinkman Lab is a state of the art facility devoted to manufacturing, and CfD uses it for precision machining, e.g., using a CNC machine. The Center for Electronics Manufacturing and Assembly ([CEMA](#)) focuses on electronics packaging, including semiconductor chip packaging, printed circuit board assemblies and electronics/optoelectronics. The CfD utilizes CEMA to populate surface mount components onto printed circuit boards (PCBs) that are designed in house. The facilities at CEMA are also ideal for reworking PCBs that need to be repaired or that require modifications.



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Cryogenic multi-layer circuit board designed in the CFD and populated in CEMA. Components on this board will be exposed to temperatures as low as 40 K, nanoTorr pressure, and high energy particle radiation.

Many other RIT facilities are available for use by the Center for Detectors. The Brinkman Lab is a state of the art facility devoted to manufacturing, and CFD uses it for precision machining, e.g., using a CNC machine. The Center for Electronics Manufacturing and Assembly (CEMA) focuses on electronics packaging, including semiconductor chip packaging, printed circuit board assemblies and electronics/optoelectronics. The CFD utilizes CEMA to populate surface mount components onto printed circuit boards (PCBs) that are designed in house. The facilities at CEMA are also ideal for reworking PCBs that need to be repaired or that require modifications.

In addition to fabrication and testing capabilities, the Center for Detectors has access to sophisticated simulation software to predict the performance of devices, from fabrication processes to performance of a completed device. Silvaco Athena and Atlas are powerful software engines that simulate the effects of processing on device substrates and the electrical characteristics of a fabricated device. Athena simulations can describe all of the processes available in the RIT SMFL, building a physics-based model in 3D space of a device from initial substrate to completed device. From this model, the user can determine doping profiles, carrier densities, depletion widths, junction depths, and dielectric layer thickness, among other parameters, all from initial conditions set by actual fabrication process recipes. Multiple devices can be simulated in a single model, so both inter- and intra-pixel performance can be characterized. Atlas can then simulate performance by deriving, e.g., current-voltage, photovoltaic, charge distribution, point spread function, and quantum efficiency (both internal and external) through user-defined operating conditions.



CfD Projects

Cosmic Radiation Damaged Image Repair project (CRDIR)

RIT has embarked on the Cosmic Radiation Damaged Image Repair project (CRDIR) involving Dr. Donald Pettit, NASA Astronaut & International Space Station astrophotographer. This project is being by graduate and undergraduate student researchers under the guidance of Dr. Donald Figer, Director of CfD.

Students will have a unique opportunity to communicate with an active American astronaut in the process of solving a significant image processing issue. The scientific process these students will follow toward a solution will provide valuable experience that these students will carry with them in their professional lives.

RIT will be providing a sophisticated "image enhancement" software program which specifically addresses degraded images taken by astronauts "on orbit", extending the useful lives of cameras, and in many cases making unacceptable "noisy" images visually acceptable.

Characterization of Inter-Pixel Capacitive Coupling in Hybridized HgCdTe Arrays

Inter-Pixel Capacitance (IPC) is a mechanism for deterministic electronic cross talk that results from coupling fields between adjacent pixels as a signal is collected and stored. Simulation of small arrays from first principles using software which simultaneously solves Poisson's equation and the Drift Diffusion equations allows for characterization of this coupling across a broad range of design parameters as well as across various environment parameters. Due to the deterministic nature of this cross-talk characterization results in correction.

This project is currently working on characterization of HgCdTe arrays hybridized using indium bumps to H2RG readout circuits akin to those to be used in the James Webb Space Telescope's (JWST) NIRcam. Successful characterization across environment parameters will result in an increase in final image quality from JWST's NIRcam and any device using a similar detector while also introducing new design considerations for future generations of sensor.

Development of Si-MOSFET CMOS Technology for Terahertz Detection

RIT is developing a silicon MOSFET CMOS imager to detect terahertz (THz) frequencies in a collaboration with the Center for Emerging and Innovative Sciences (CEIS) at the University of Rochester and Exelis Geospatial Systems. Creating an asymmetrical design within the FETs, increases the THz response. The current device being tested was designed with 15 individual test transistors with varying design dimensions and antennas along with an array of transistors for an imager. These test structures are being evaluated at RIT to determine the best design for terahertz detection for future imager designs. An advantage of this detector technology is that these MOSFETs will not have to be cooled to extreme temperatures like microbolometers. This project will advance knowledge of the detection mechanism, lead to the creation of an integrated imaging system, and has many different applications.

THz frequencies have been largely unexplored due to high absorption within water, however there has been an increase in interest with the rise in high-altitude and space-based telescopes. Emission lines in spectra within the THz regime exhibit cool molecular gas which traces protoplanetary disks and star formation rates within galaxies. This technology will also have other applications within the medical and security fields because of the non ionizing, non-harmful nature of THz radiation.



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Phase II: New Infrared Detectors for Astrophysics

This program will have a profound impact across ground-based and space-based astronomy by dramatically reducing the cost of infrared detectors for existing facilities, as well as the next generation of extremely large telescopes. The project will continue development of a new material system for use in astronomical infrared array detectors. The devices use HgCdTe grown on Silicon using Molecular Beam Epitaxy. In Phase I of this project, the NSF ATI program funded two cycles of design, fabrication, and testing. The devices made in this phase of the project show that the technology will be successful for astronomy requirements pending further development. Testing shows that there are several challenges that prevent the devices in hand from satisfying these requirements. In this second phase, the effort will develop a series of devices with improved design and processing. The approach is to reduce the number of material defects while maintaining high short-wavelength quantum efficiency by using proven designs. Given that these previous designs have successfully been used to address the observed non-idealities, it is believed that the new activities will be successful. The work includes: growth of new material using a thick buffer layer design, fabrication of twelve FPAs in two designs, and extensive testing between the fabrication of the two designs.

This project will advance the knowledge of a material system that has great promise for infrared detector technology. It will enhance the capabilities of infrared instrumentation in astronomy by reducing cost and potentially improving performance when compared to what is available with existing technology. The technical approach has great merit because it was developed over the past 15 years and during Phase I of the project. The plan features a tight connection between design at Raytheon Vision Systems (RVS) and testing in the Center for Detectors at the Rochester Institute of Technology, continuing over 15 years of collaboration between RVS and the PI.

Single Photon Counting Detectors for NASA Astronomy Missions

Single photon counting detectors have the potential to be the next big advancement for NASA astronomy missions. The ability to count single photons facilitates science goals that are impossible even with current state-of-the-art detectors. Single photon counting detectors are the future, and many different implementations are in development. In the next 20 years, many NASA missions requiring single photon counting will be proposed, but which single photon counting detector implementations best suit the performance needs of NASA's astronomy programs? The goal of the proposed research is to characterize (theoretically and physically) three unique implementations of single photon counting detectors, benchmark their operation over a range of performance characteristics, and provide comprehensive justification for the superiority of one of the implementations for each of these NASA astronomy applications: exoplanet detection, high-contrast imaging, adaptive optics, and array-based LIDAR.

The research plan in this project includes simulation, characterization, and evaluation of the performance of three types of semiconductor photon counting detectors for use in NASA astronomy missions: Geiger-mode (GM) APDs, linear-mode (LM) APDs, and Electron Multiplying (EM) CCDs.

The main goal of this project is to provide the basis of comparisons for several types of fundamentally different single photon counting detectors by producing a table of comparisons and recommendations for various applications. All detectors must be sensitive to single photons, be scalable to large array formats, and have high QE (in the visible, UV, NIR, IR, and Far-IR wavelengths). This research will advance preliminary work by adding new characterization methods and performance benchmarks, and by comparing the different devices at predetermined milestones during the project. The recommended detector(s) should function well at high readout frequencies without significant read noise (leading to improved temporal sampling), have very low noise for increased SNR at low fluence levels, and be implemented (or able to be in the near future) on large arrays. Even though single-photon counting detectors share a common



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sampling), have very low noise for increased SNR at low fluence levels, and be implemented (or able to be in the near future) on large arrays. Even though single-photon counting detectors share a common performance benchmark (discerning individual quanta), differences between various implementations make some more efficient than others.

New Infrared Detectors for Astrophysics

Infrared arrays with HgCdTe as the light-sensitive layer, such as have been developed up to sizes 2048x2048 pixels for the James Webb Space Telescope, are near-ideal detectors for imaging and spectroscopy in the region ~1-5 microns. However current construction requires fabrication on CdZnTe substrates, which are expensive and limited in availability. The key to making larger (up to 14,000x14,000 pixels) and less expensive infrared detectors lies in using silicon wafer substrates, since large silicon wafers are common in the high volume semiconductor industry and their coefficient of thermal expansion is well-matched to that of the silicon readout circuits.

While the use of silicon substrates has been a major goal in the field of developing infrared detectors, the main limitation over the past 15 years has been the large lattice spacing mismatch between silicon and commonly-used infrared light-sensitive materials. The mismatch causes defects that can result in higher dark current, or valence holes that lead to reduced quantum efficiency and image persistence.

Enlisting the expertise and fabrication capabilities of Raytheon Vision Systems, detector expert Dr. D. Figer of the Rochester Institute of Technology plans to deposit the HgCdTe light-sensitive layer on silicon using the very promising technique of Molecular Beam Epitaxy (MBE). By maintaining vacuum during MBE processing, defect density has been shown to be reduced and the resulting prototype devices have achieved the anticipated performance. Very large, affordable infrared arrays will be essential for making optimum use of the proposed ~30m class ground-based telescopes and their availability has clear implications for fields beyond astronomy, including medical imaging and remote sensing.

A Zero Read Noise Detector for Thirty Meter Telescope (TMT)

The key objective of this project is to develop a new type of imaging detector that will enable the most sensitive possible observations with the world's largest telescopes, i.e. the TMT. The detector will effectively quadruple the collecting power of the TMT, compared to detectors currently envisioned in TMT instrument studies, for the lowest light level observations. It would have fundamental importance in ground-based and space-based astrophysics, Earth and planetary remote sensing, exo-planet identification, consumer imaging applications, and homeland safety, among many others. Measurable outcomes include being able to see further back into the infancy of the Universe to taking a better picture (less grainy) of a smiling child blowing out the candles at her birthday party. The detector will be quantum-limited (zero read noise), be resilient against the harsh effects of radiation in space, consume low power, operate over an extremely high dynamic range, and be able to operate with exposure times over one million times faster than typical digital cameras. The RIDL is teaming with MIT/Lincoln Laboratory to leverage their Geiger-mode Avalanche Photodiode technology for developing the imaging detector. The project is funded through the [Gordon and Betty Moore Foundation](#).

A Photon Counting Detector for Exoplanet Missions

The objective of this project is to advance photon-counting detectors for NASA exoplanet missions. A photon counting detector will provide zero read noise, ultra-high dynamic range, and ideal linearity over the relevant flux range of interest. It could be the best realizable detector for a planet finding spectrograph, and it would have outstanding properties as a wavefront sensor or detector in the imaging focal plane. The benefit of this device is that it will dramatically improve sensitivity beyond what is capable with non-photon



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it would have outstanding properties as a wavefront sensor or detector in the imaging focal plane. The benefit of this device is that it will dramatically improve sensitivity beyond what is capable with non-photon counting detectors, thereby increasing science return for a fixed mission life; for low-light-level cases, such as spectroscopy, the device will reduce the necessary exposure time for detecting planetary features by 50-80%. The device always operates in photon counting mode and is thus not susceptible to excess noise factor that afflicts other technologies. It continues operating with shot-noise limited performance up to extremely high flux levels. Its performance is expected to be maintained at a high level throughout mission lifetime in the presence of the expected radiation dose.

Past Projects

LIDAR Imaging Detector

In collaboration with MIT/Lincoln Laboratory, the RIDL is developing an imaging Light Detection and Ranging (LIDAR) detector for NASA planetary space missions. The device will have a pixelated array of independent Geiger-mode Avalanche Photodiodes that can asynchronously measure laser light time of flight. The output will be three-dimensional images providing distance measurements for each pixel. The device will have a timing accuracy of ~100 picoseconds, thus enabling a ranging accuracy ~1 cm, or roughly two orders of magnitude better than existing LIDAR instruments.

The nature of GLIMPSE 81: a star cluster to rival Westerlund 1

This project uses Chandra/ACIS observations of a young star cluster. The X-ray emission from this cluster, already observed in previous low-resolution observations, will be resolved by Chandra into many components. Analysis will include separation of the diffuse X-ray emission from the point-sources, and a spectral analysis of each source. The data from this project will be combined with those from other observations in order to perform a multiwavelength analysis.

Clumping in OB-star winds

Massive stars, their nature and evolution, play an important role at all stages of the Universe. Through their radiatively driven winds they influence on the dynamics and energetics of the interstellar medium. The winds of OB stars are the most studied case. Commonly, the mass-loss rates of luminous OB stars are inferred from several types of measurements, the strengths of UV P Cygni lines, H-alpha emission and radio and FIR continuum emission. Recent evidence indicates that currently accepted mass-loss rates may need to be revised downwards when small-scale density inhomogeneities (clumping) are taken into account. This project uses Herschel Space Telescope data to consistently treat ALL possible diagnostics, scanning different parts of the winds, and analyzed by means of 'state of the art' model atmospheres, will permit the determination of true mass-loss rates.

Guide Detector Evaluation for the Large Synoptic Survey Telescope (LSST)

The RIDL is testing prototype optical guide detectors for LSST. We have developed a guider testbed that simulates the expected image motion that the LSST focal plane will experience. The testbed will be used to validate guide detector performance against LSST requirements.

Sigma-Delta Detector

In collaboration with the University of Rochester, the RIDL is developing very low noise CMOS imaging detectors for NASA space astrophysics and planetary missions. These devices promise sub-electron read noise in a direct-readout architecture that is resilient against the transient and long-term effects of radiation.



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detectors for NASA space astrophysics and planetary missions. These devices promise sub-electron read noise in a direct-readout architecture that is resilient against the transient and long-term effects of radiation. The novel readout circuit uses a one-bit sigma-delta oversampling comparator design developed by Zeljko Ignjatovic and Mark Bocko (University of Rochester). The light-sensitive silicon wafer is being designed and fabricated by RIDL and the [RIT Semiconductor & Microsystems Fabrication Laboratory](#). NASA/JPL will delta-dope the backside of the detector for enhanced ultraviolet sensitivity. The wafer and the readout will be bump bonded together to produce a hybrid detector with good sensitivity from the ultraviolet through near-optical infrared.

A NICMOS survey of newly-identified young massive clusters

We are on the cusp of a revolution in massive star research triggered by 2MASS and Spitzer/GLIMPSE, and now is the ideal time to capitalize on these projects by performing the first survey of massive stars in young stellar clusters throughout the Galactic plane. A search of the 2MASS and GLIMPSE surveys has produced over 450 newly-identified massive stellar cluster candidates in the Galactic plane which are hidden from our view at optical wavelengths due to extinction. In this project, we are using 29 HST orbits to image the most promising candidate clusters in broad and narrow band filters using NICMOS. The observations will be complemented with approved Spitzer and Chandra programs, numerous approved and planned groundbased spectroscopic observations, and state-of-the-art modelling. We expect to substantially increase the numbers of massive stars known in the Galaxy, including main sequence OB stars and post-main sequence stars in the Red Supergiant, Luminous Blue Variable and Wolf-Rayet stages. Ultimately, this program will address many of the fundamental topics in astrophysics: the slope to the initial mass function (IMF), an upper limit to the masses of stars, the formation and evolution of the most massive stars, gamma-ray burst (GRB) progenitors, the chemical enrichment of the interstellar medium, and nature of the first stars in the Universe.

Mid-Infrared Spectroscopy of the Most Massive Stars

The most massive star that can form is presently defined by observations of a class of very rare stars having inferred initial masses of \sim 200 solar masses. There are only a few such stars in the Galaxy, including the Pistol Star, FMM362, and LBV 1806-20, the first two being located near the Galactic center, and third located in the disk near W31. Each has only recently been identified as so massive within the past 10 years through the analysis of infrared observations, but they are otherwise too faint, due to extinction, to observe at shorter wavelengths. These stars appear to be very luminous ($L > 10^{6.3}$ solar luminosities), "blue" ($T > 10000$ K), and variable ($\Delta K \sim 1$ mag.), and the Pistol Star has ejected 10 solar masses of material in the past 10000 years. In addition, these stars have near-infrared spectra similar to those of prototypical Luminous Blue Variables, i.e. Eta Car and AG Car. Given their apparent violation of the Humphries-Davidson limit, they are presumably in a short-lived phase of stellar evolution that is often associated with rapid mass-loss through episodic eruptions of their outer atmospheres. We propose to determine the physical properties of these stars and the velocity and ionization structure in their winds by using spectra obtained with the high resolution modes of the Infrared Spectrograph (IRS) on the Spitzer Space Telescope. The 10 to 40 micron wavelength region is ideally suited for accessing a variety of lines from transitions of hydrogen, helium, iron, silicon, sulfur, among others; indeed, through our models, we predict that sufficiently sensitive spectra will yield over 300 spectral lines. In addition, we predict that the mid-infrared continuum will be dominated by free-free emission generated in the thick winds associated with these stars, an effect that should be clearly detectable in the spectra.

Sandia FPGA Imaging Acquisition Software Development

RIDL developed HDL (Hardware Description Language) software for a FPGA (Field Programmable Gate Array) evaluation board for Sandia National Laboratories. The HDL software programs the FPGA chip to



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Sandia FPGA Imaging Acquisition Software Development

RIDL developed HDL (Hardware Description Language) software for a FPGA (Field Programmable Gate Array) evaluation board for Sandia National Laboratories. The HDL software programs the FPGA chip to transfer digital video rate image data from the evaluation board to a windows based PC using the Camera Link data transfer format. Once the Camera Link data format became available, data acquisition software for the PC was developed to capture and display the imagery on the PC.

The Pre-Supernova Mass-Loss Behavior of Red Supergiants

The mass lost by massive stars as they pass through the Red Supergiant (RSG) phase is a crucial determinant in the terminal mass of the star, the nature of the resulting supernova explosion and the stellar end-state. However, up until now studies of this quantity have been problematic, owing to the low numbers of known RSGs and the difficulty of observing in the mid-IR. Here, we capitalize on the recent discoveries of two remarkable Galactic clusters containing unprecedented numbers of RSGs, and use the capabilities of Spitzer to undertake a comprehensive and unique study of the pre-SN mass-loss of massive stars. We will use Spitzer/IRS observations in conjunction with state-of-the-art dust models to provide the first quantitative investigation of the mass and composition of the pre-SN ejecta as a function of age and metallicity. This study is vital in determining the mass-loss behaviour of RSGs in particular, and the nature of supernova progenitors in general.

A Very Low Noise CMOS Detector Design for NASA

The purpose of this project is to design, fabricate, and test a novel new detector having many advantages over present CCD and CMOS devices. The hybrid CMOS detector is expected to have low noise, radiation tolerance, low power dissipation, low mass and robust electronics. Its low noise is derived from a novel on-die digitization circuit based on a sigma-delta feedback loop.

The Most Massive Stars

Until the Spitzer Space Telescope, there was no wide area survey that could identify massive stars at all distances in the Galaxy. Indeed, the sample of known O-stars is woefully incomplete, as it has largely been generated using optical observations that suffer from the absorption produced by dust in the disk. We will find and measure the physical properties of the most massive stars in the Galaxy using HST, Spitzer, Chandra, SOFIA, and ground-based observatories, using a survey technique that probes the majority of the Galaxy. This program addresses fundamental questions whose answers are basic requirements for studying many of the most important topics in Astrophysics: the formation and evolution of the most massive stars, the effects of massive stars on lower mass protostellar/protoplanetary systems, gamma-ray burst progenitors, nature of the first stars in the Universe, chemical enrichment of the interstellar medium, Galactic gas dynamics, star formation in starbursts and merging galaxies (particularly in the early Universe). The results of our program will influence the science programs for future NASA projects, i.e. JWST, SOFIA, SIM, TPF-C, and TPC-I.



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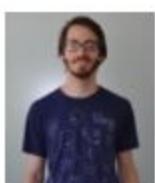
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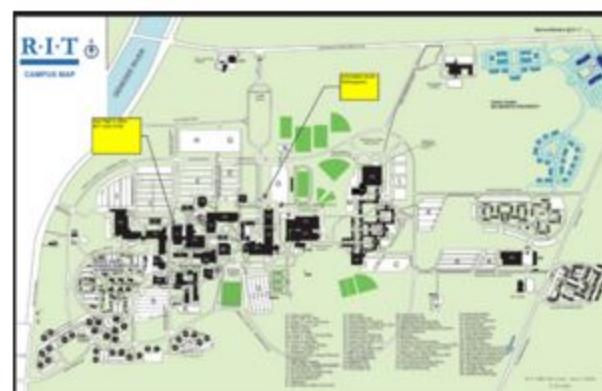
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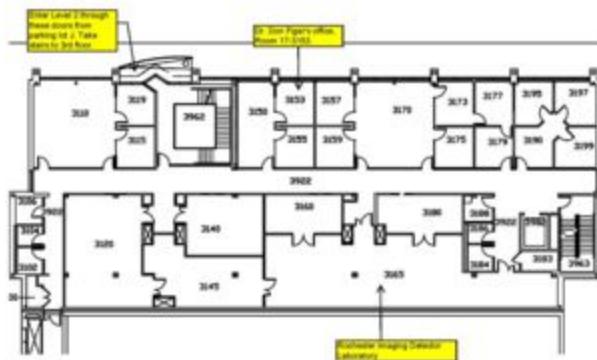
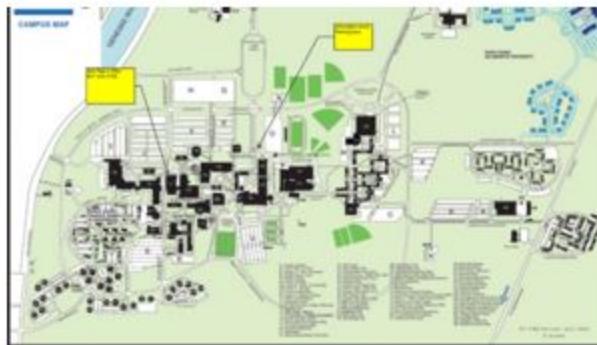
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Friday, April 10, 2015 9:43 AM	<dir> other labs
Friday, March 11, 2016 1:54 PM	<dir> projects
Wednesday, June 03, 2015 3:13 PM	<dir> ridlcam
Monday, October 26, 2015 10:30 AM	<dir> software



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